

The 17th CNS International Summer School



MEAN-FIELD STUDY OF THE RADIATIVE CAPTURE $^{12}\text{C}(p,\gamma)^{13}\text{N}$ AND $^{13}\text{C}(p,\gamma)^{14}\text{N}$ REACTIONS

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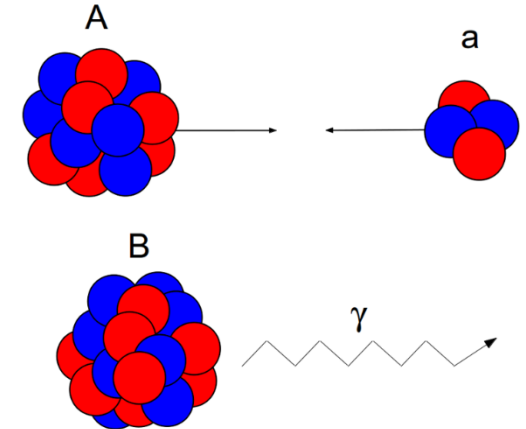
CONTENT

- I. Radiative capture
- II. Nuclear mean-field potential
- III. Mean-field description of the $^{12}\text{C}(p,\gamma)^{13}\text{N}$ and $^{13}\text{C}(p,\gamma)^{14}\text{N}$ reactions
- IV. Summary

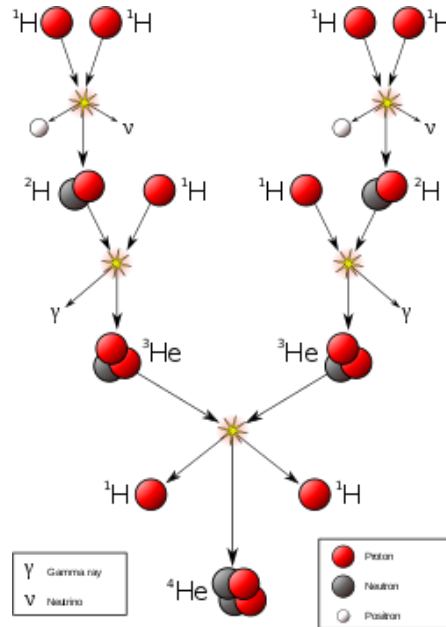
RADIATIVE CAPTURE

Radiative capture is an important process due to its astrophysical applications.

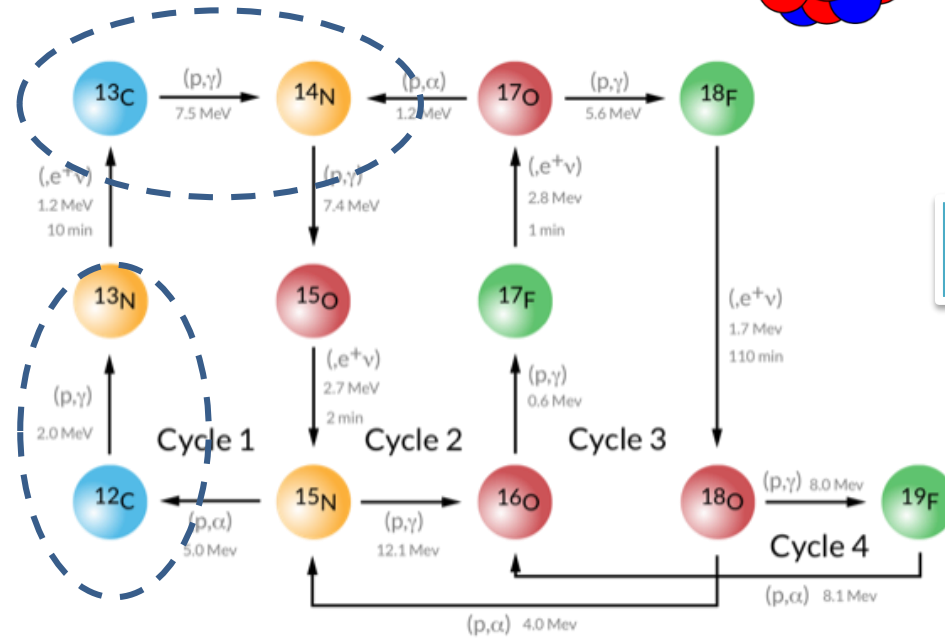
BBN, stellar evolution, element synthesis, X-ray bursts, etc.



pp chain



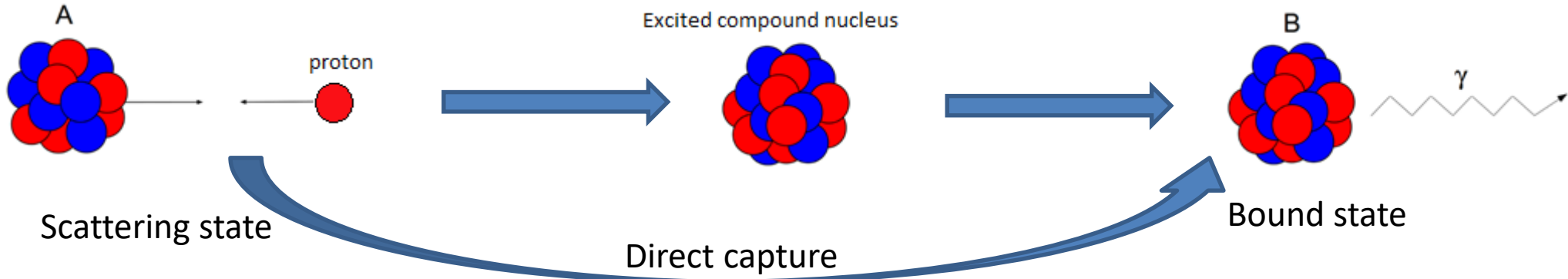
Hydrogen burning



CNO cycle

CNO: $T_9 < 0.2$

RADIATIVE CAPTURE



The radial Schrödinger equation:

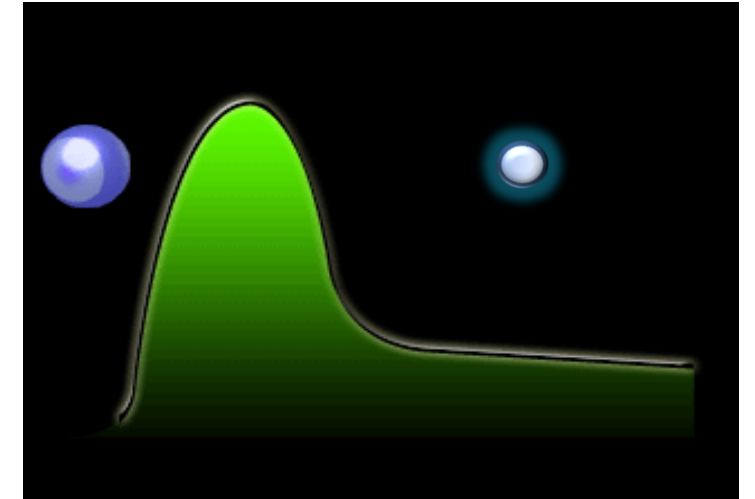
$$\frac{\hbar^2}{2\mu} \left[\frac{d^2}{dr^2} - \frac{l(l+1)}{r^2} \right] u(r) + V(r)u(r) = Eu(r)$$

{ Bound state
} Scattering state

Coulomb pot. Nuclear pot. Spin-orbit pot.

Normalization

{	Bound state:	$u_j(r) \rightarrow C \exp(-k_B r)$
}	Scattering state:	$u_j(r) \rightarrow F_j(kr) \cos \delta_j + G_j(kr) \sin \delta_j$



RADIATIVE CAPTURE

Using the balanced detail, the cross section for the radiative capture $A(p,\gamma)B$ reaction is determined as

$$\sigma(J_f, E) \sim \sum_{\sigma\lambda} \underbrace{\left(\frac{E_\gamma}{\hbar c}\right)^{2\lambda-1}}_{\text{Photon}} \underbrace{|\langle \Psi^{J_f} \| M^{\sigma\lambda} \| \Psi(E) \rangle|^2}_{\text{Nucleon}}$$

Matrix elements needed for electromagnetic transitions

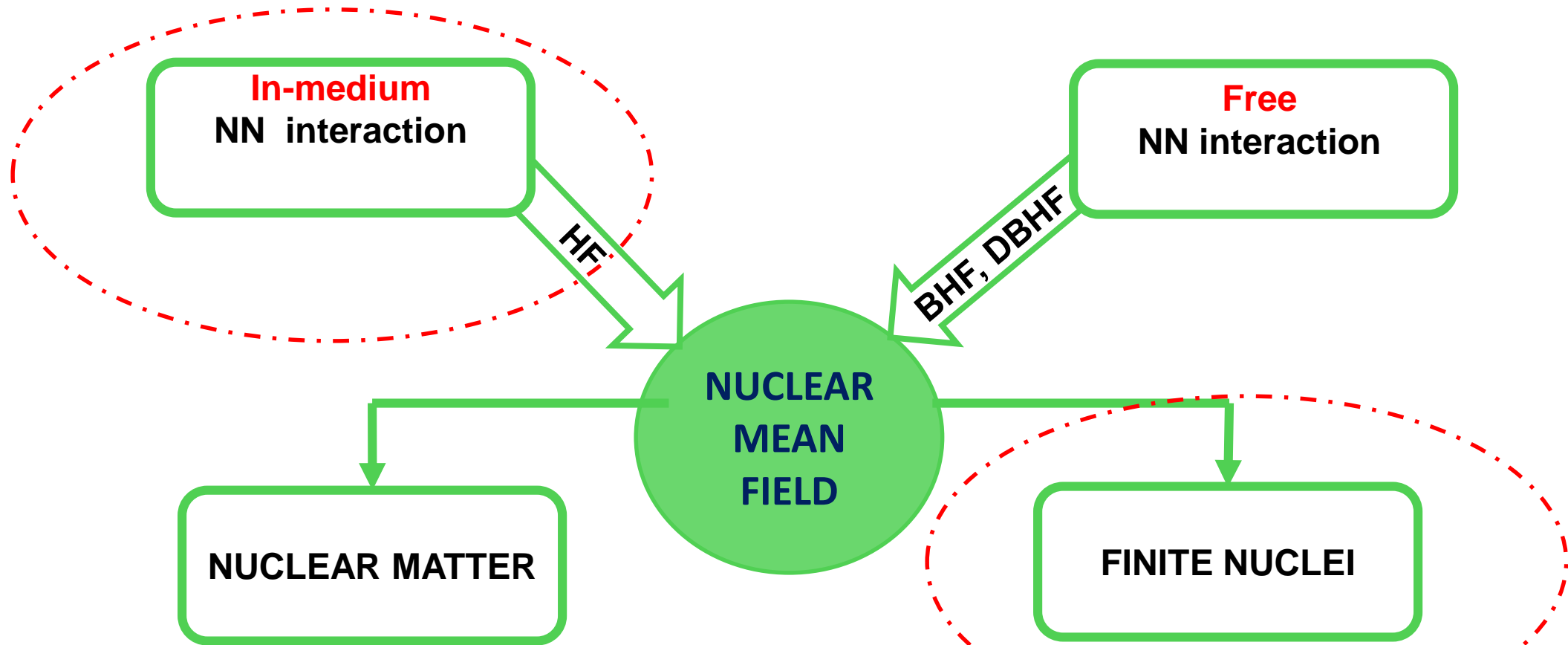
$$\langle \Psi^{J_f} \| M^{\sigma\lambda} \| \Psi(E) \rangle = eZ_{eff} \underbrace{\langle J_f \lambda M_f \mu | J_i M_i \rangle}_{\text{Coupling coefficient}} \sqrt{\frac{(2J_i + 1)(2\lambda + 1)}{4\pi(2J_f + 1)}} \int_0^\infty \underbrace{u_i(E, r) r^\lambda u_f(r) dr}_{\text{Overlap integral}}$$

Scattering state
Bound state

Long wavelength approximation

$$\sigma\lambda = \underbrace{E_1}_{\text{Photon}} \gg E_2 \approx M_1 \gg E_3 \approx M_2, \dots$$

NUCLEAR MEAN-FIELD POTENTIAL



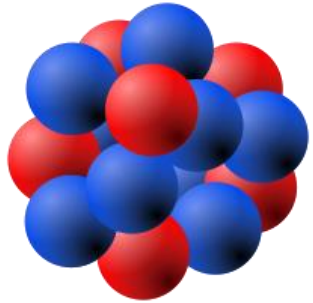
- ✓ EOS of nuclear matter
- ✓ Single particle potential
- ✓ Symmetry energy
- ✓ Neutron star

Ngo Hai Tan *et al*, *Phys. Rev. C* 93, 035806
Doan Thi Loan *et al*, *Phys. Rev. C* 92, 034304
Dao T. Khoa *et al*, *Phys. Rev. C* 94, 034612

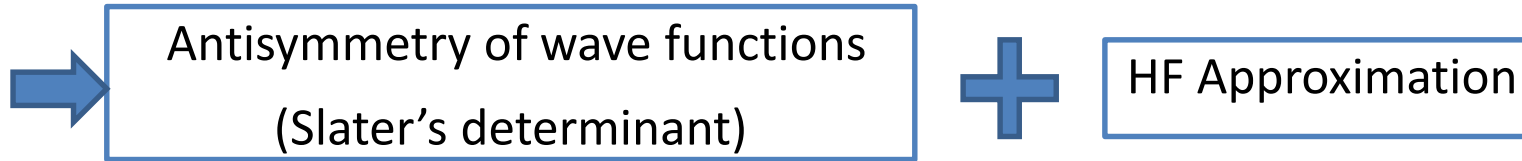
- ✓ Bound problems
- ✓ Scattering problems

Radiative capture

NUCLEAR MEAN-FIELD POTENTIAL



Fermions



$$v = v^D + v^{EX}$$

M3Y

In-medium (density dependent) NN interaction

$$v^{D(EX)}(\rho, s) = F_0(\rho)v_{00}^{D(EX)}(s) + F_1(\rho)v_{01}^{D(EX)}(s)\vec{\tau}_1 \cdot \vec{\tau}_2 \quad \text{with } s = |\vec{r}_1 - \vec{r}_2|$$

Density dependence

D.T. Khoa, G.R. Satchler and W. von Oertzen, *Phys. Rev. C* 56, 954 (1997);
D.T. Loan, B.M. Loc, and D.T. Khoa, *Phys. Rev. C* 92, 034304 (2015).

G-matrix based on M3Y interaction

N. Anantaraman, H. Toki, G.F. Bertsch, *Nucl. Phys. A* 398, 269 (1983).

HF calculation

HvH theorem

Extended HF (HF + RT) calculation

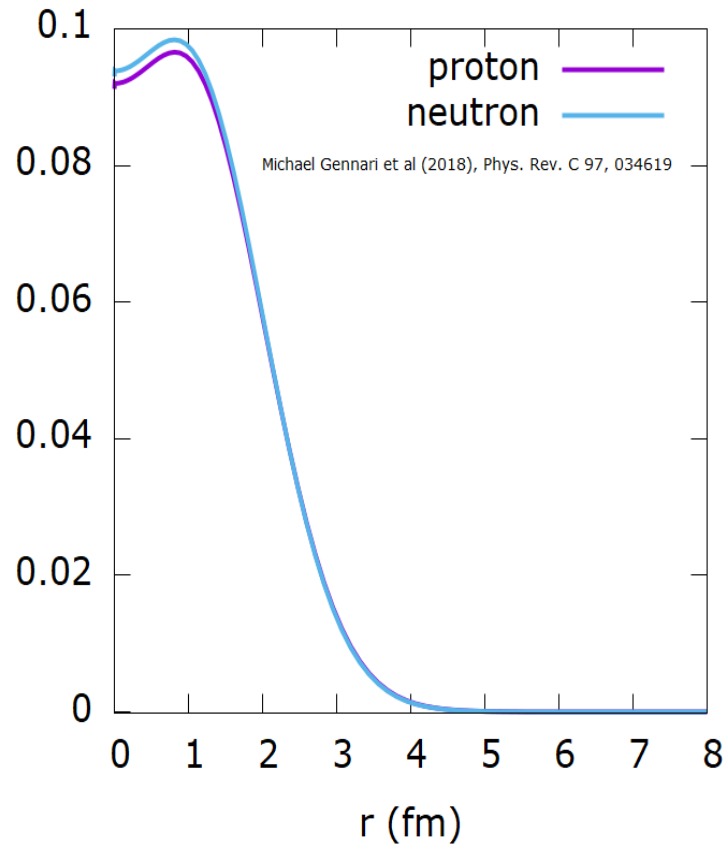
Doan Thi Loan, Bui Minh Loc, and Dao T. Khoa, *Phys. Rev. C* 92, 034304 (2015)

Dao T. Khoa, Nguyen Hoang Phuc, Doan Thi Loan, and Bui Minh Loc, *Phys. Rev. C* 94, 034612 (2016)

NUCLEAR MEAN-FIELD POTENTIAL

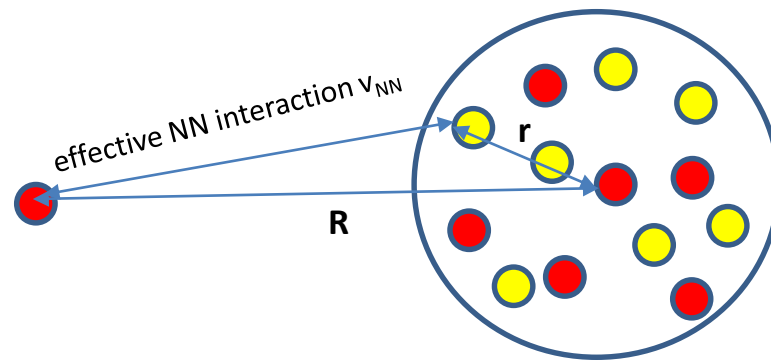
Single folding model

Nucleon density of ^{12}C at ground state



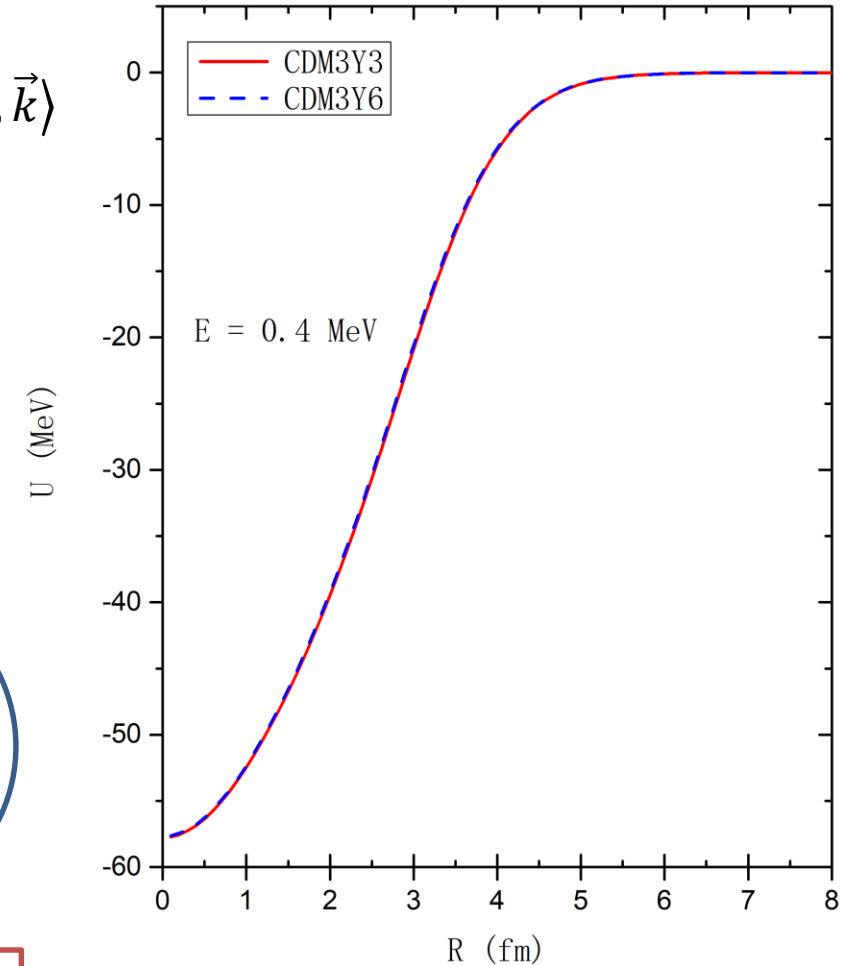
$$U = \sum_{j \in A} \langle \vec{k}, j | v_c^D | \vec{k}, j \rangle + \langle \vec{k}, j | v_c^{EX} | j, \vec{k} \rangle$$

$$U(\vec{R}) = \int dr \rho_A(\vec{r}) v_{NN}(\vec{r} - \vec{R}, \rho)$$



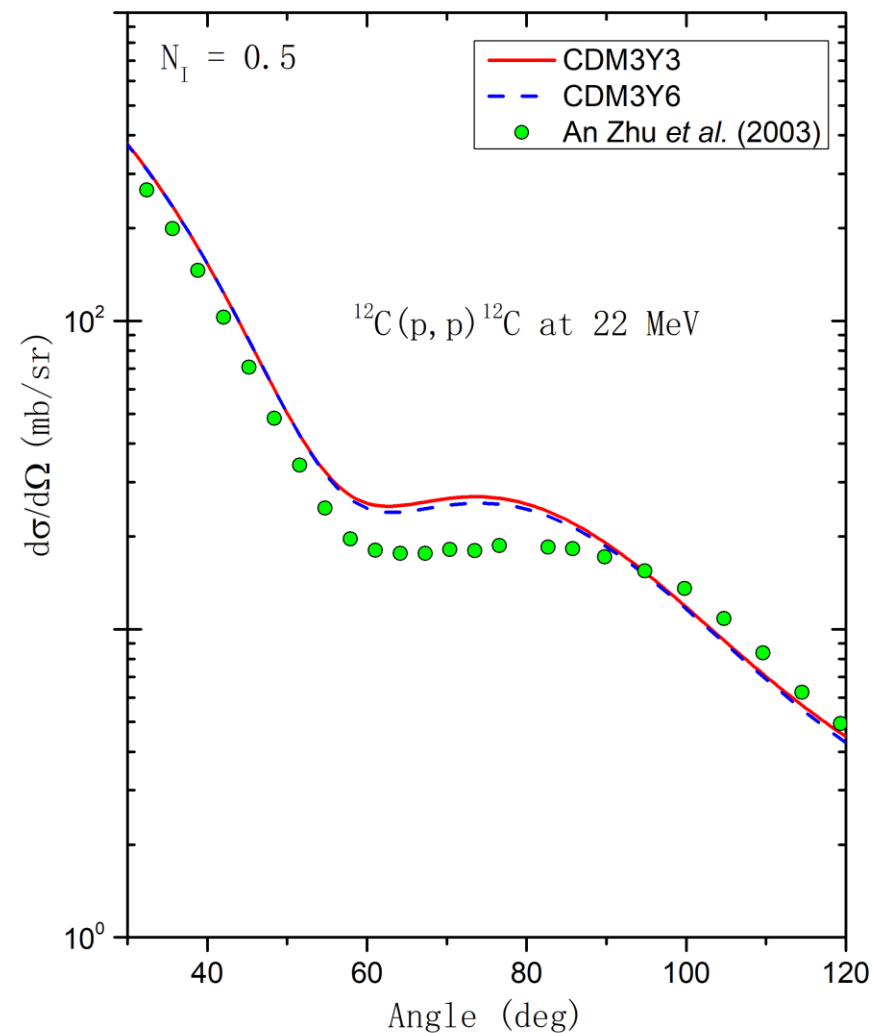
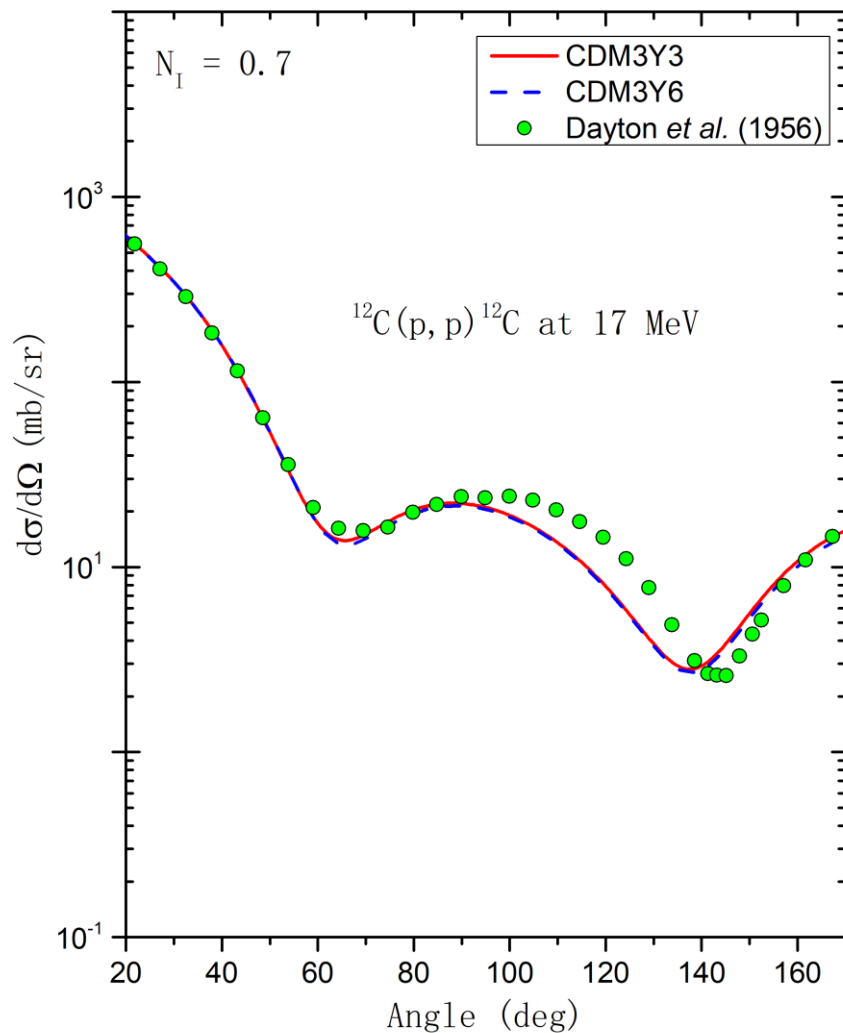
mean-field

Folding model potential of p + ^{12}C scattering

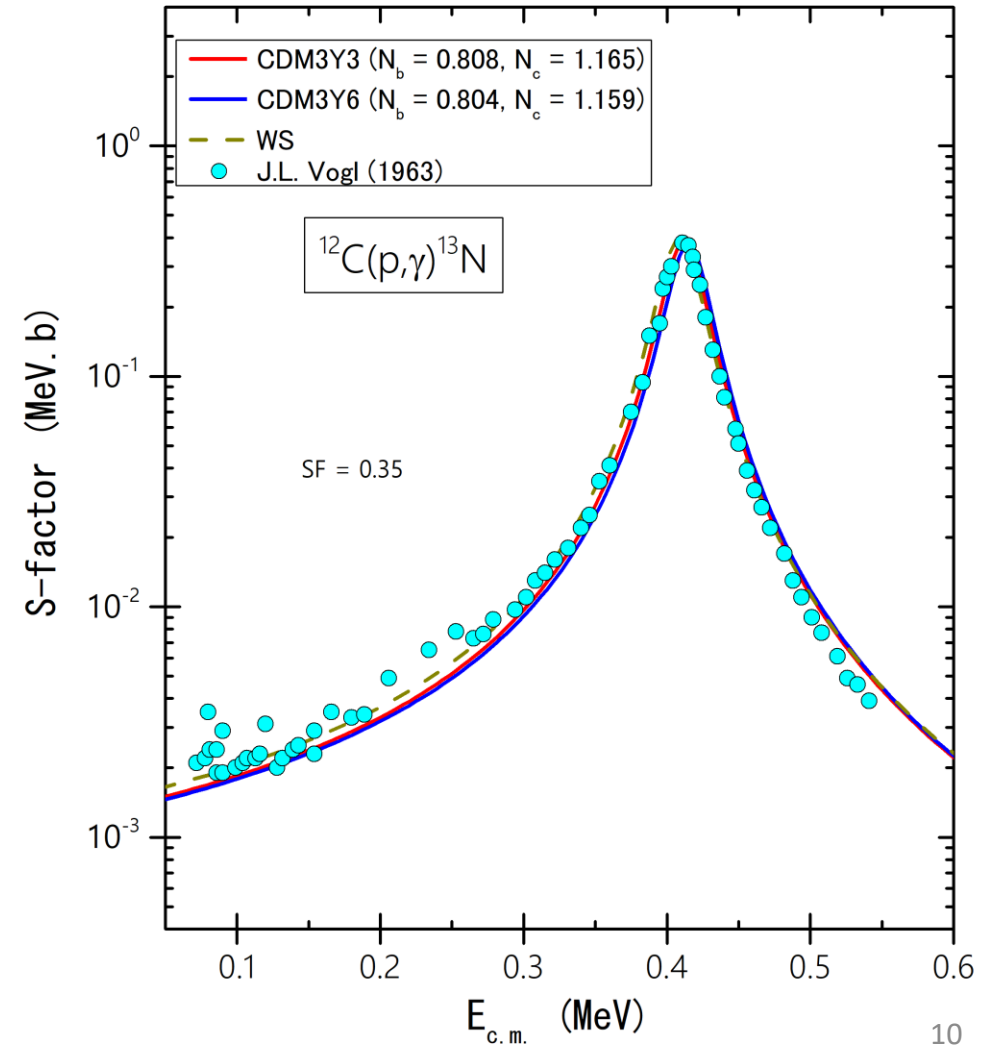
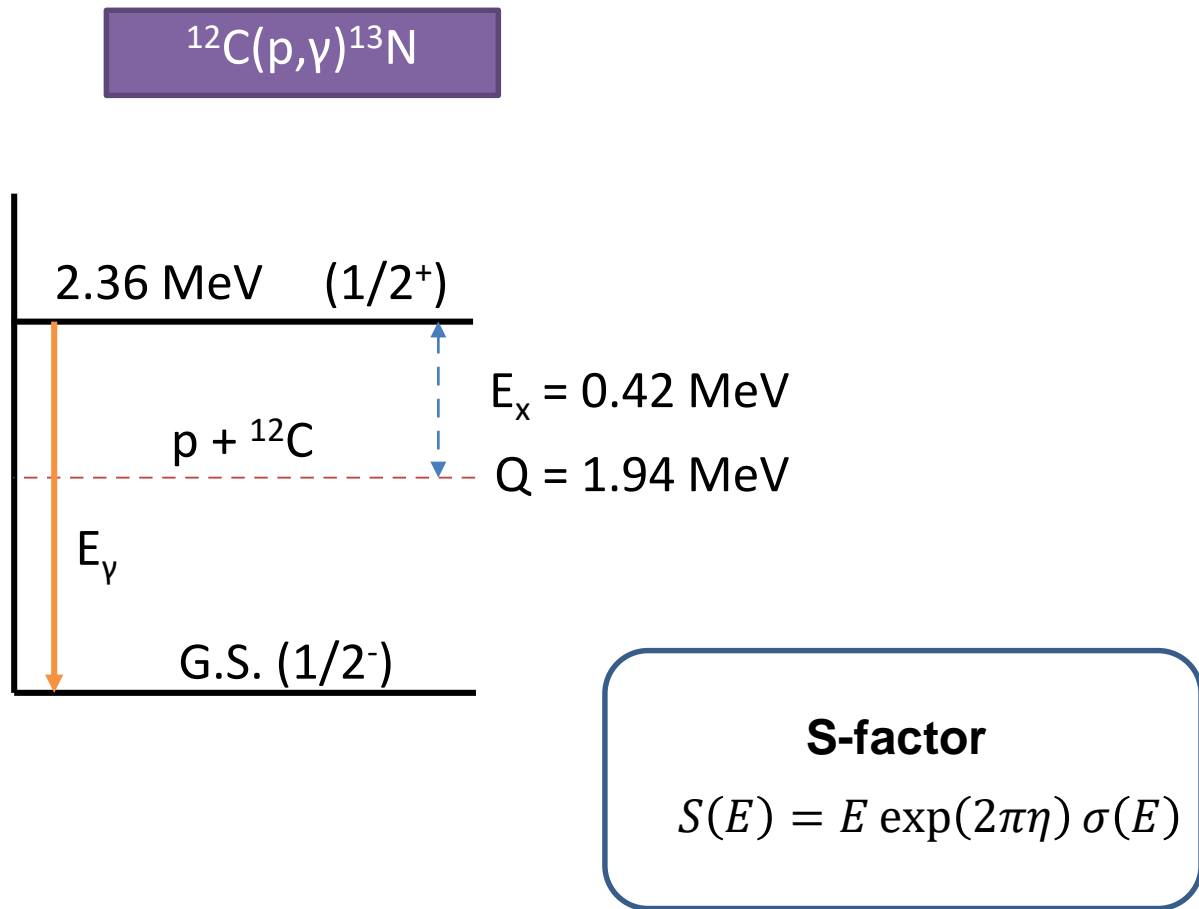


NUCLEAR MEAN-FIELD POTENTIAL

Elastic (p + ^{12}C) scattering



MEAN-FIELD DESCRIPTION OF THE $^{12}\text{C}(p,\gamma)^{13}\text{N}$ AND $^{13}\text{C}(p,\gamma)^{14}\text{N}$ REACTIONS

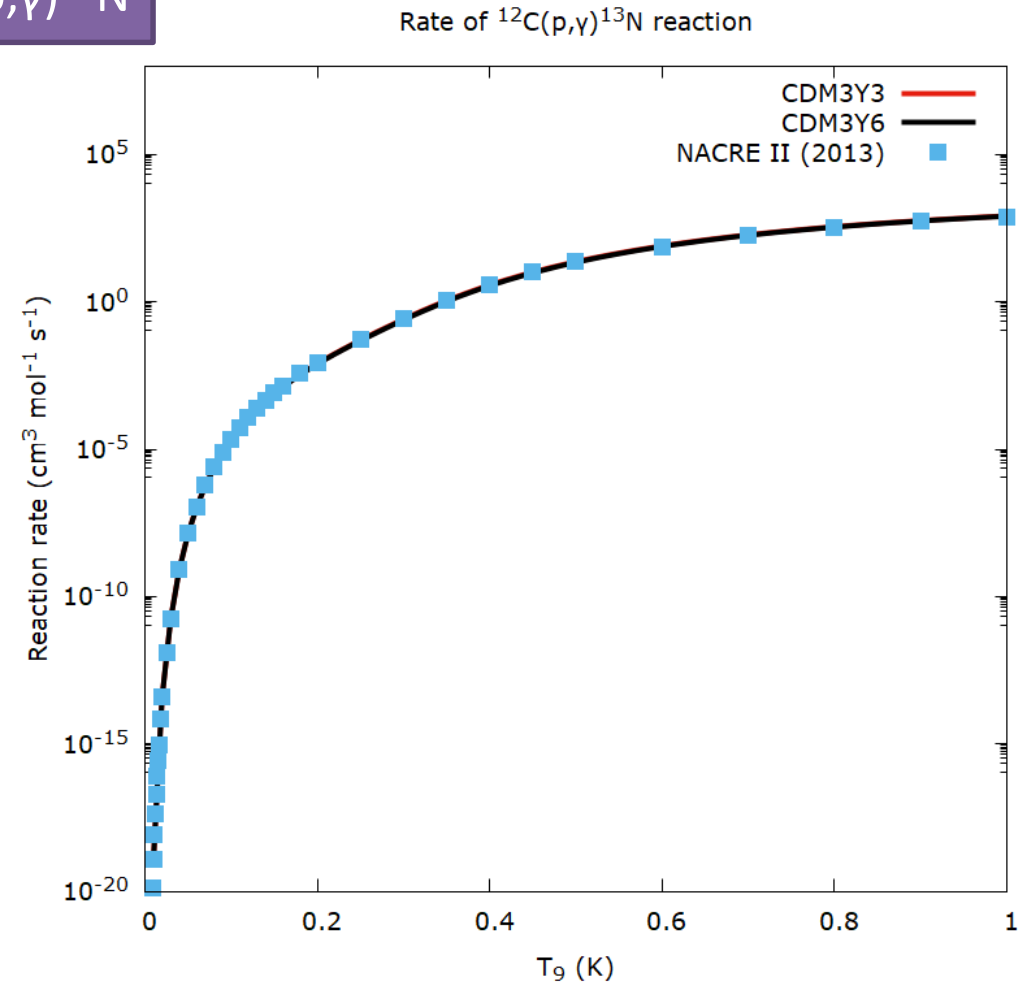


MEAN-FIELD DESCRIPTION OF THE $^{12}\text{C}(p,\gamma)^{13}\text{N}$ AND $^{13}\text{C}(p,\gamma)^{14}\text{N}$ REACTIONS

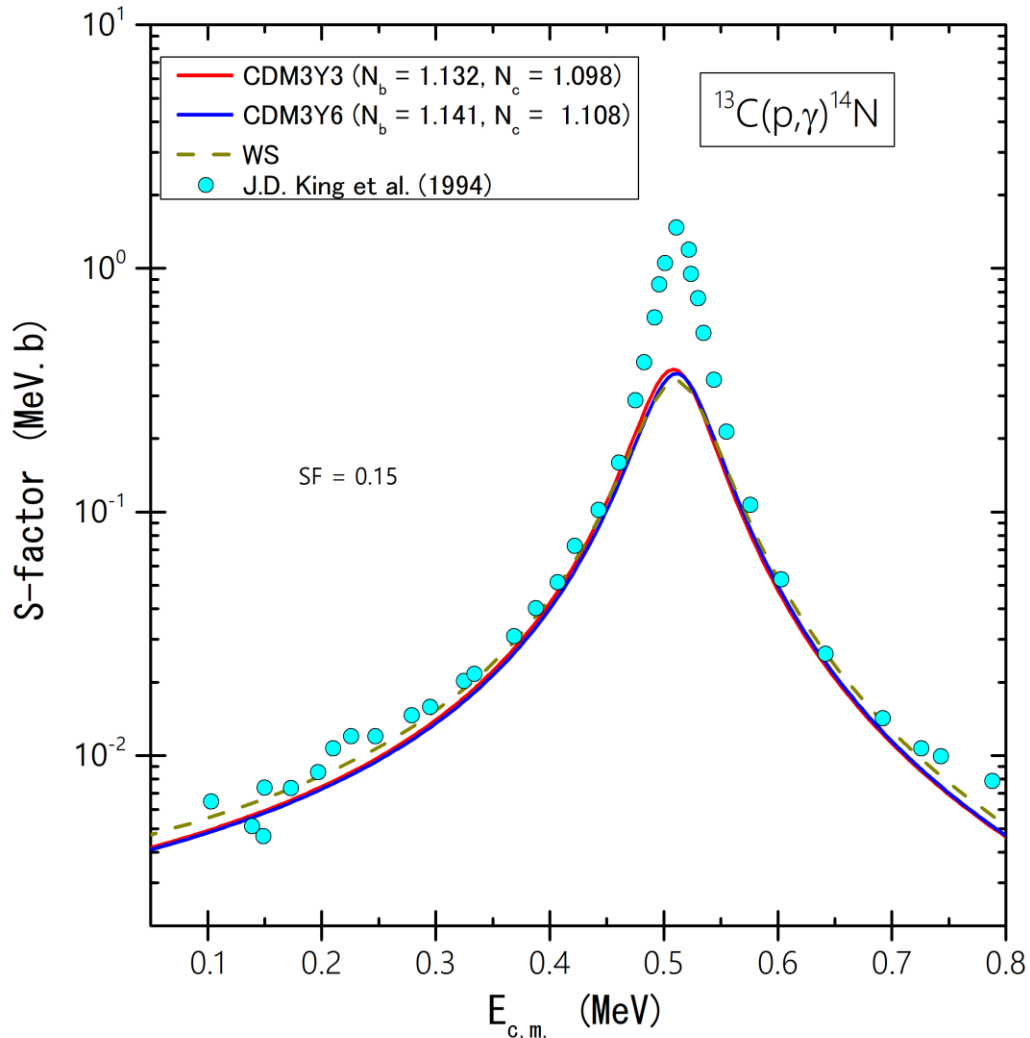
Reaction rate

$$\langle\sigma v\rangle = \left(\frac{8}{\pi\mu}\right)^{\frac{1}{2}} \frac{1}{(k_B T)^{\frac{3}{2}}} \int_0^\infty \sigma(E) E \exp\left(-\frac{E}{k_B T}\right) dE$$

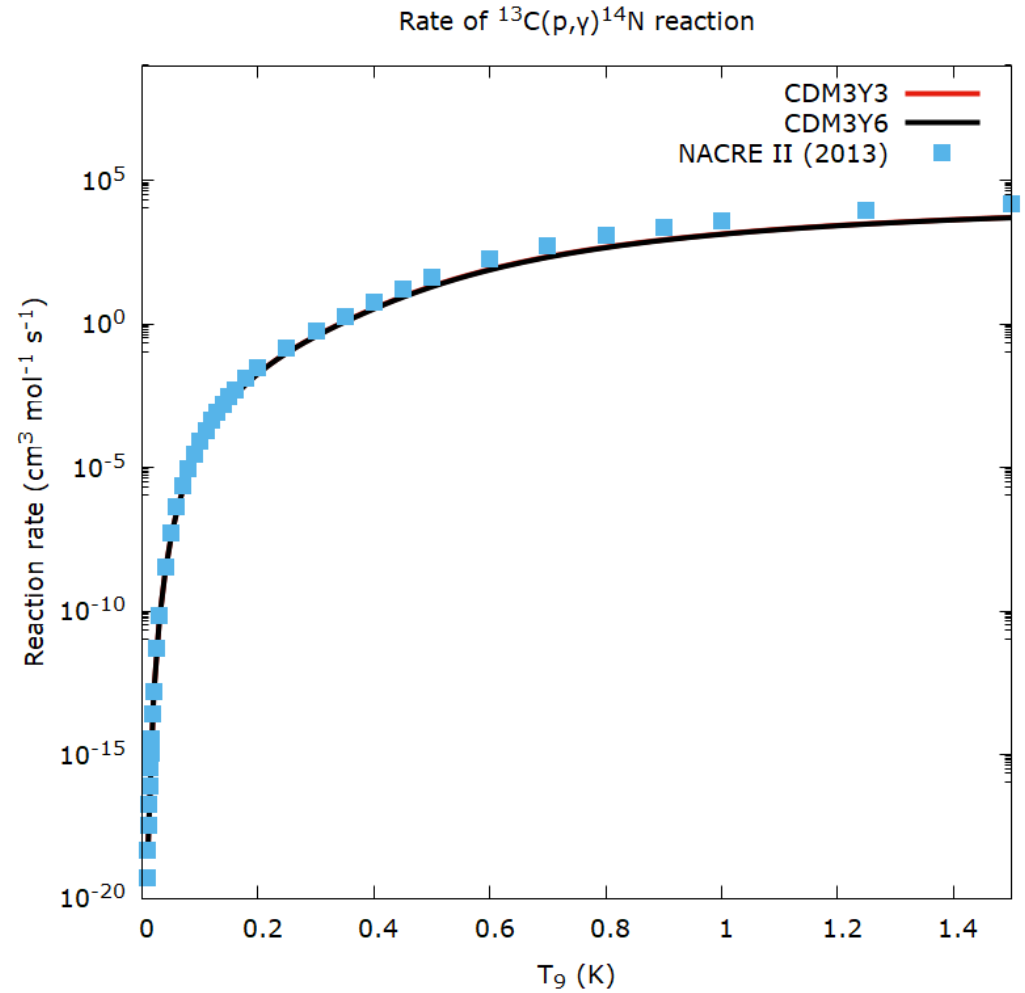
$^{12}\text{C}(p,\gamma)^{13}\text{N}$



MEAN-FIELD DESCRIPTION OF THE $^{12}\text{C}(p,\gamma)^{13}\text{N}$ AND $^{13}\text{C}(p,\gamma)^{14}\text{N}$ REACTIONS



$^{13}\text{C}(p,\gamma)^{14}\text{N}$



SUMMARY

The folded potential gives a good OM description of the elastic $p+^{12}\text{C}$ scattering at several energies.

This SFM approach is further used to calculate the nuclear mean-field potential for the study of the astrophysical S factor of the $^{12}\text{C}(p,\gamma)^{13}\text{N}$ and $^{13}\text{C}(p,\gamma)^{14}\text{N}$ reactions.

Reaction rates of the radiative capture reactions which are an importantly astrophysical quantity are produced to describe effectively the experimental data.

THANK YOU FOR YOUR ATTENTION!