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Direct measurement of the cross section for

$^{102}\text{Pd}(p, \gamma)^{103}\text{Ag}$ reaction in the p-process

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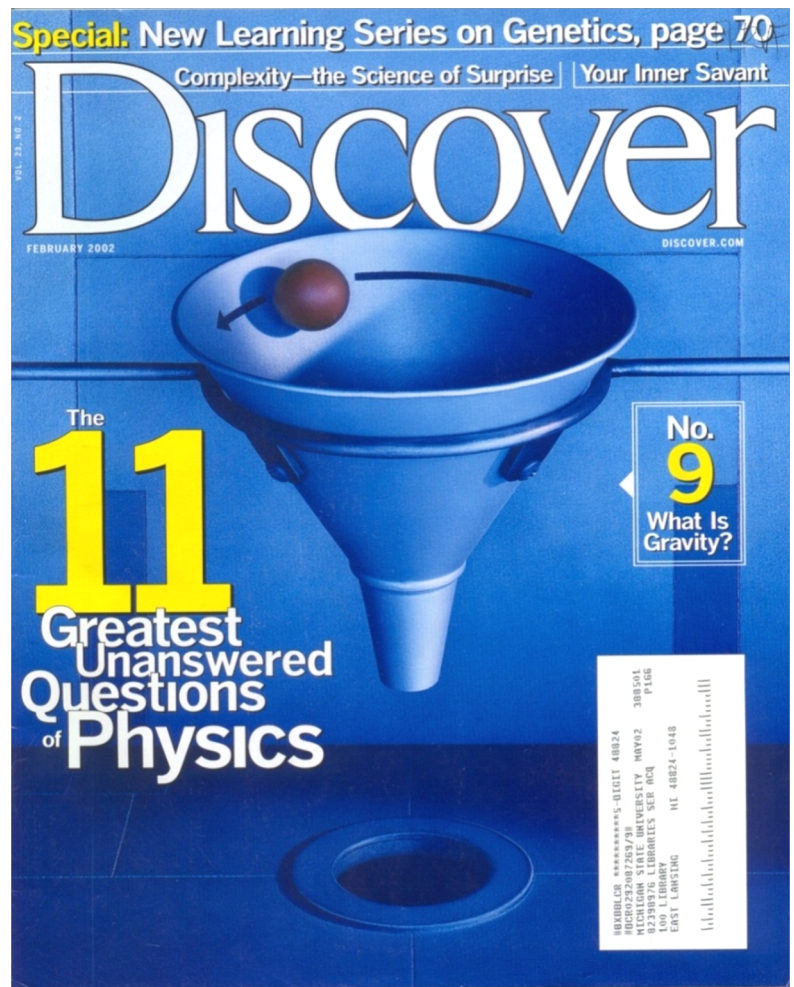
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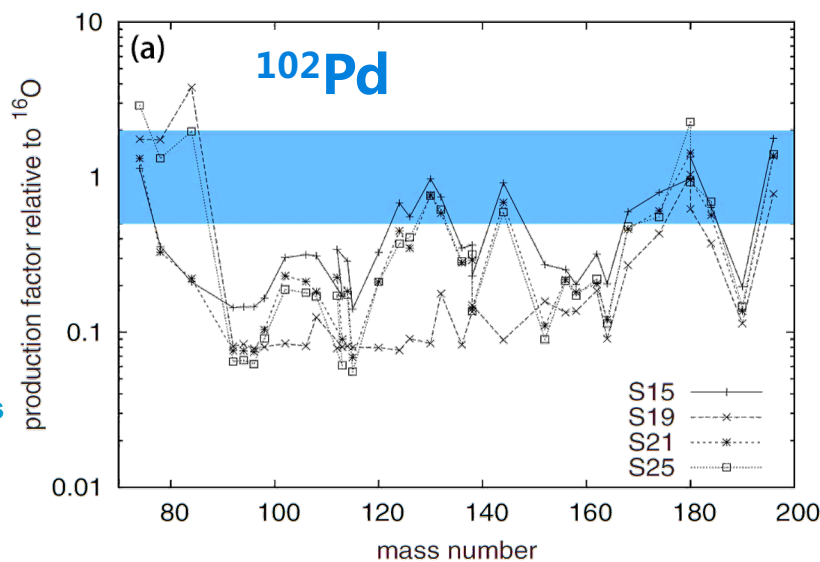
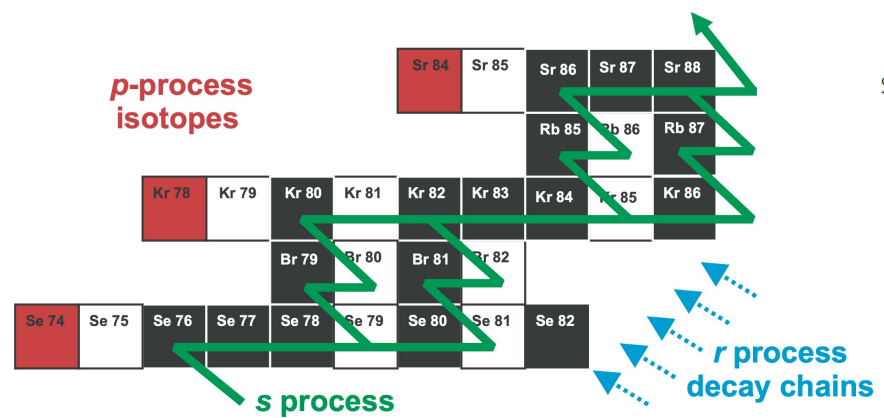
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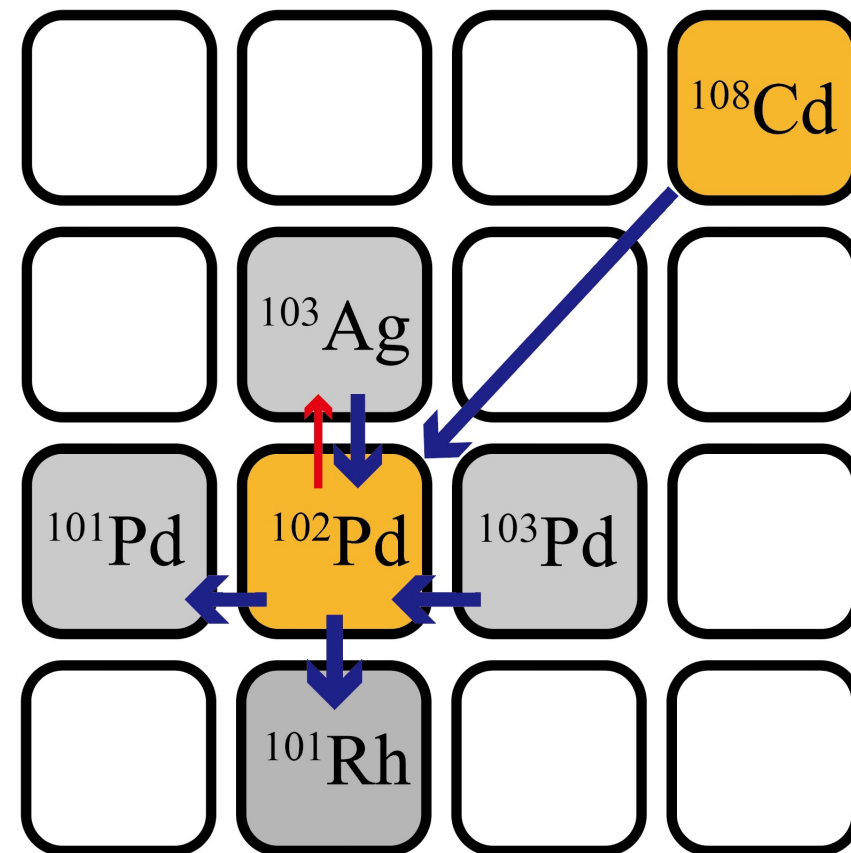
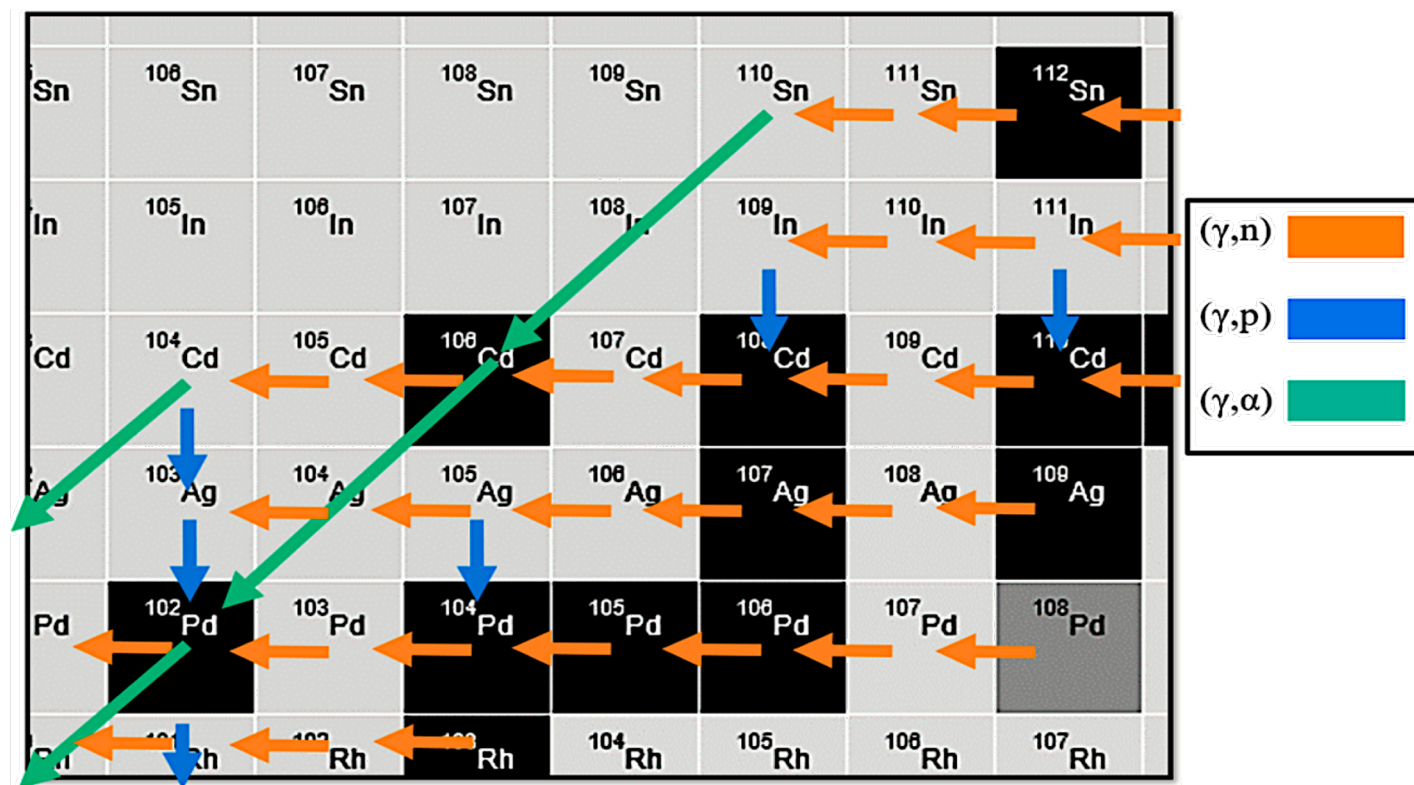
In 2002, "Discover" reported 11 unsolved mysteries of physics

1. What is dark matter?
2. What is dark energy?
3. How were the heavy elements from iron to uranium made?

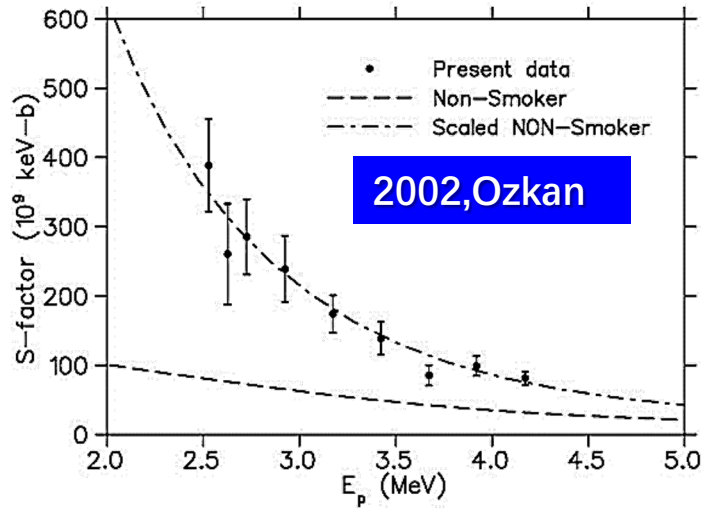


p-nuclei : Mass between ^{74}Se - ^{196}Hg , can not be synthesized through the s-process and r-process.

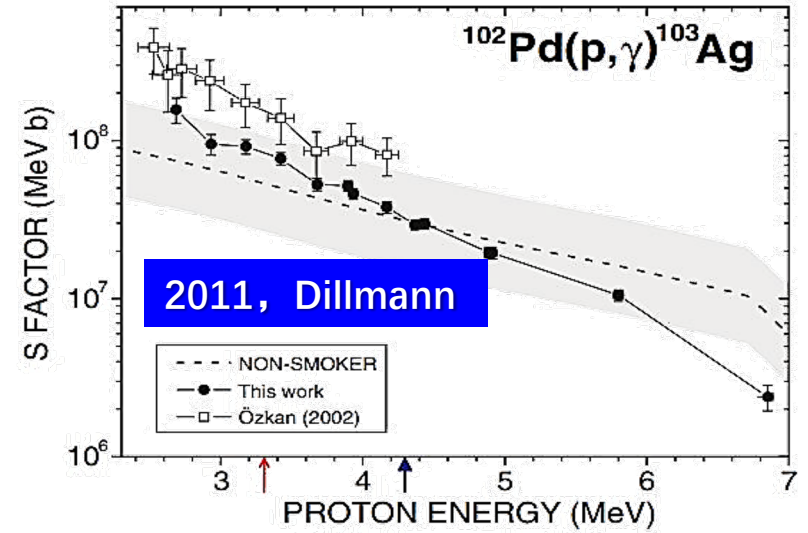
● p-production factors relative to ^{16}O , Below $A < 124$ and between $150 < A < 165$ the p-isotopes are severely underproduced.



■ The latest research shows that the reverse reaction of $^{102}\text{Pd}(p,\gamma)$ is one of the important constructive reaction of ^{102}Pd , and it is also one of the destruction reactions of ^{102}Pd

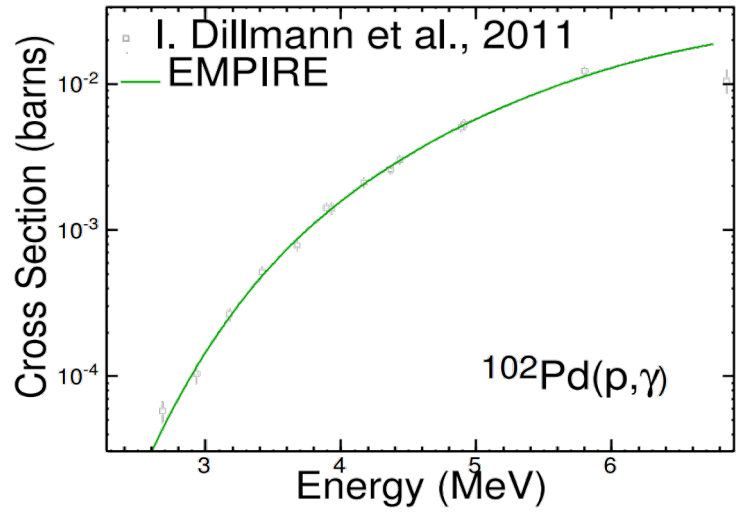
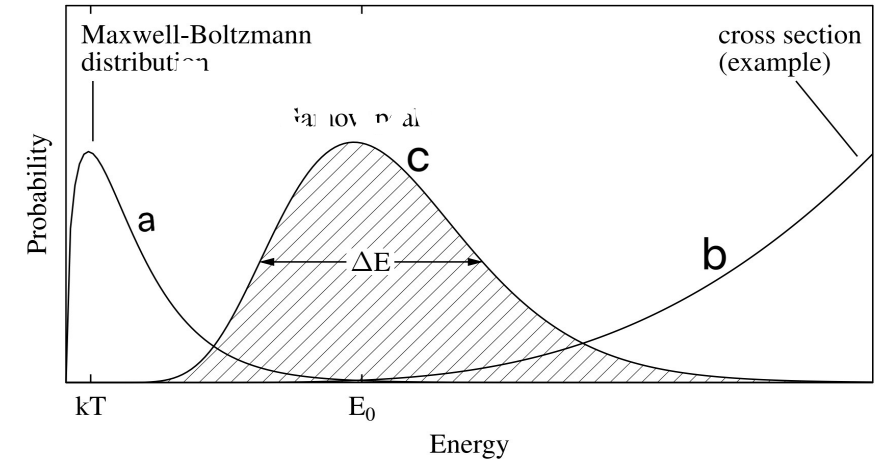


Nuclear Physics A, 2002, 710(3-4): 469-485

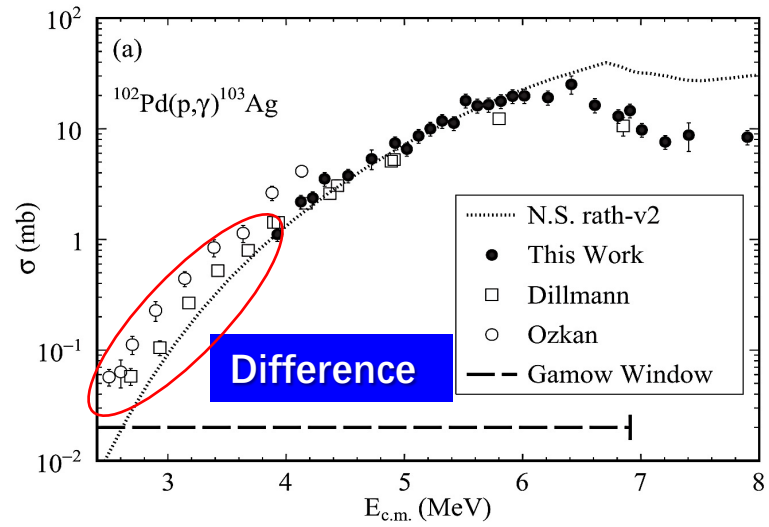


Physical Review C, 2011, 84(1): 015802

Sonnabend K., Ph. D. thesis



Nuclear Data Sheets, 2014,120: 180-183.



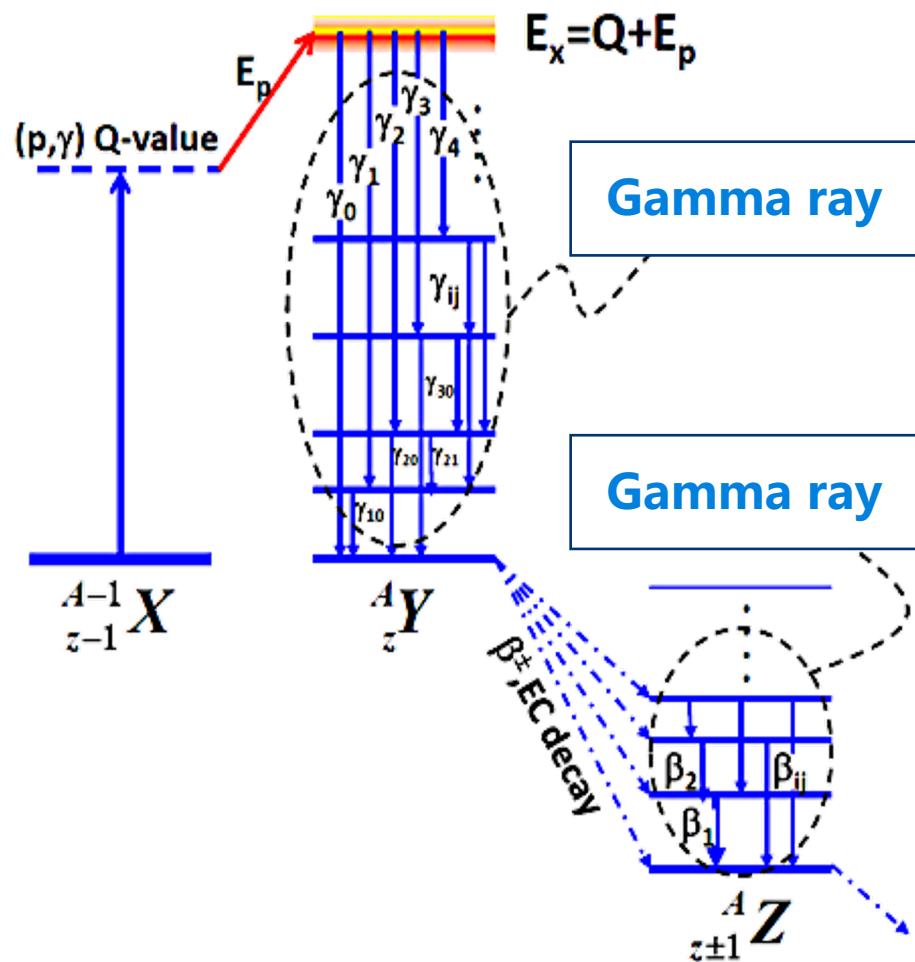
Physical Review C, 2020, 102, 055806.

◆ Experimental data for the lower energy region of greatest concern in nuclear astrophysics are lack

◆ Several times differences between the different experimental values



Experimental method :



On Line

$$\frac{dN}{dt_b} = \sigma(E)P\phi - \lambda N$$

Online section

$$N = \frac{\sigma(E)P\phi}{\lambda} (1 - e^{-\lambda t_b})$$

$$n_\gamma = N e^{-\lambda t_w} (1 - e^{-\lambda t_m}) \epsilon_\gamma \eta_\gamma$$

Offline Section

$$\sigma(E) = \frac{\lambda n_\gamma}{P\phi(1 - e^{-\lambda t_b})e^{-\lambda t_w}(1 - e^{-\lambda t_m})\epsilon_\gamma \eta_\gamma}$$

N : Number of ^{103}Ag

λ : Decay constant

t_b : Beam time

n_γ : Number of γ rays

$\sigma(E)$: Cross section

ϵ_γ : Detector efficiency

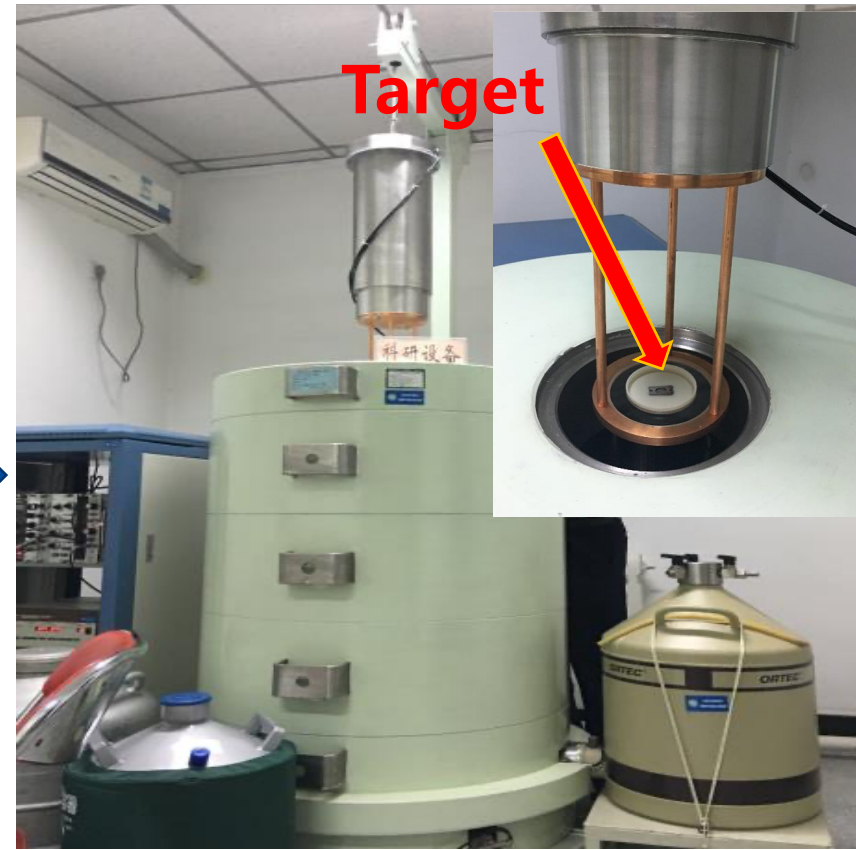
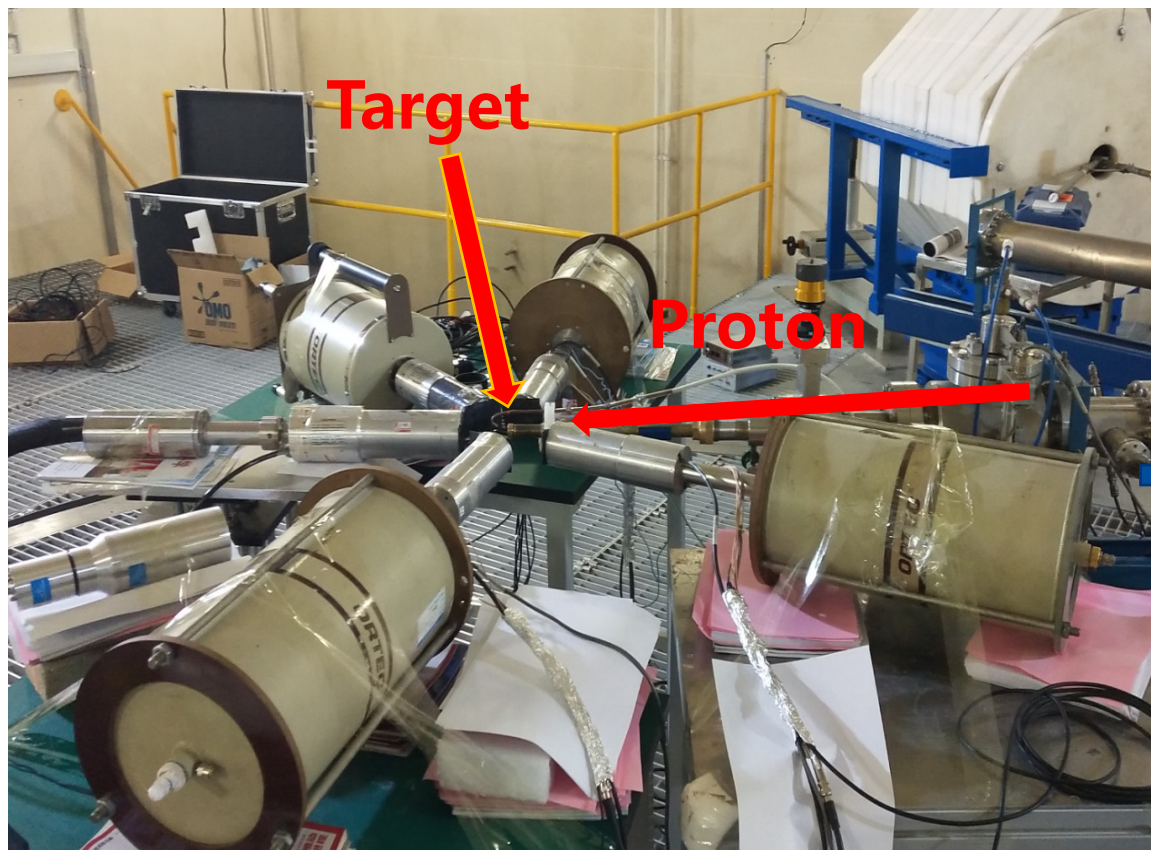
P : Density of target

$$P = \frac{T_{tar} N_A A_{Dy}}{M_{tar}}$$

η_γ : Branching ratio

ϕ : Beam intensity

Activation



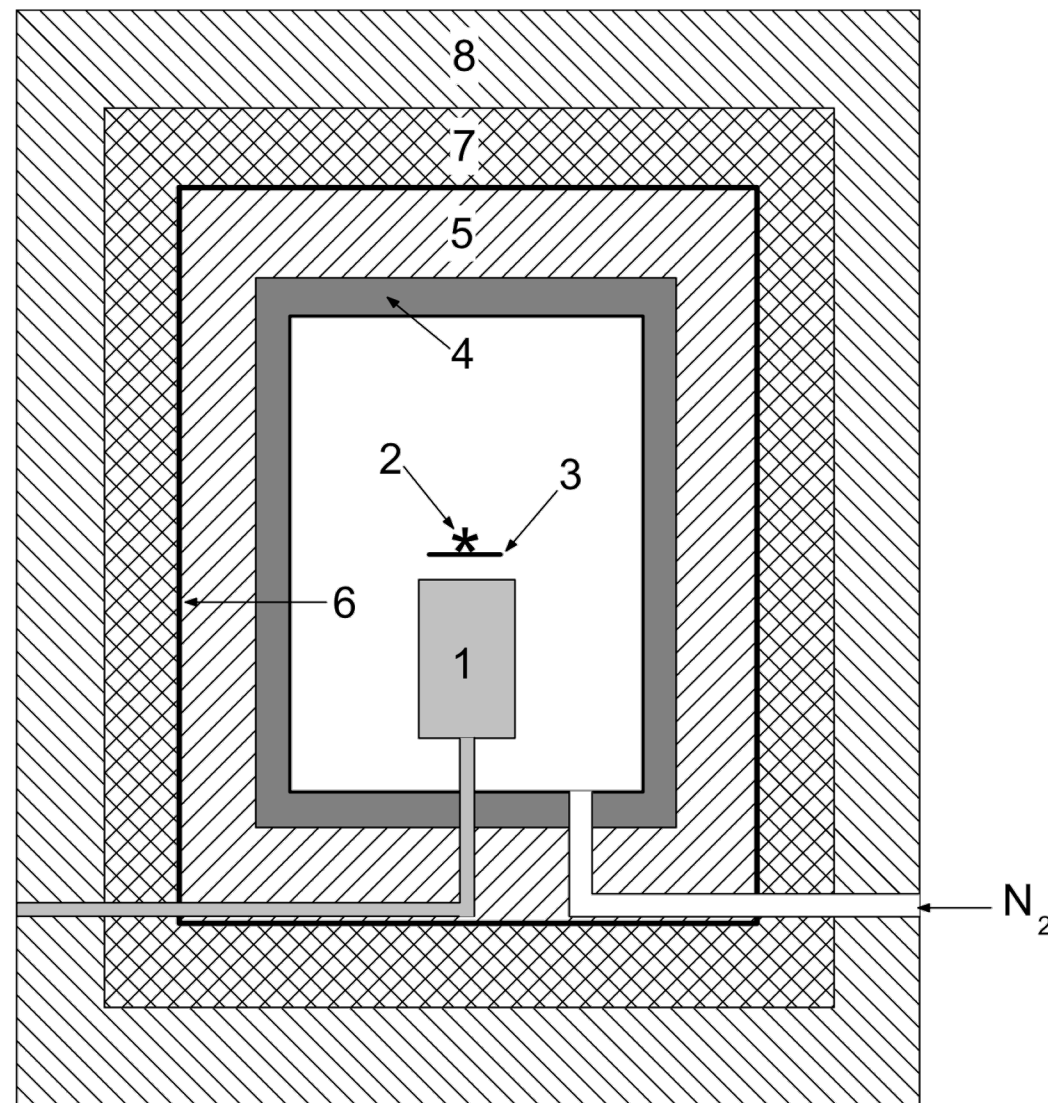
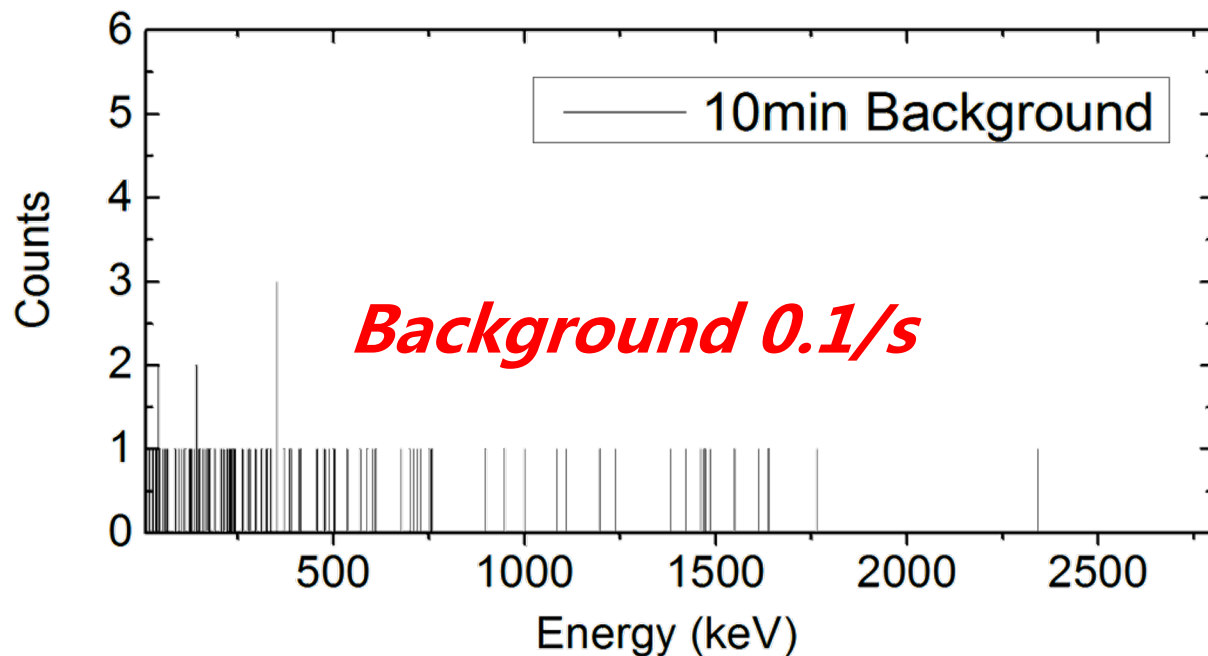
- 2×1.7 MV Tandem accelerator
- 1.9-3.2 MeV $6 \mu\text{A}$

- The low background anti-muon and anti-Compton spectrometer.

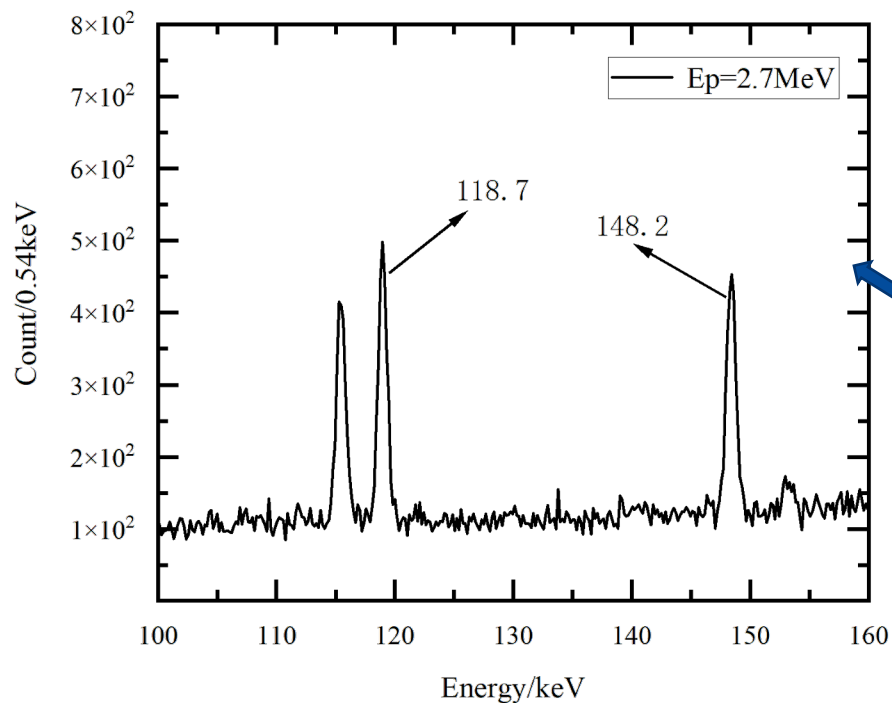


Schematic of the low-background setup

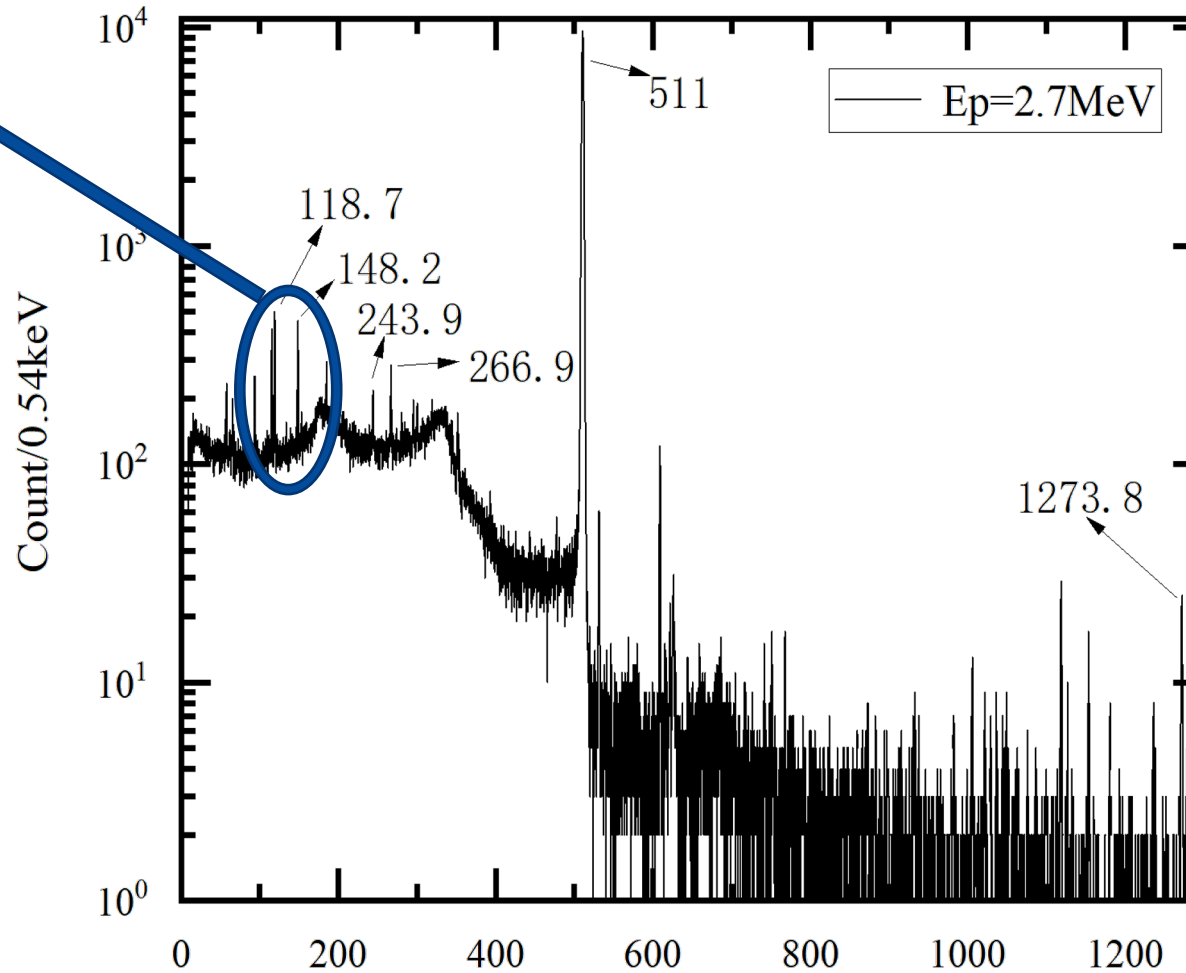
1 is the HPGe detector, 2 the source, 3 the acrylonitrile butadiene styrene (ABS) plastic holder, 4 the copper liner, 5 the inner lead ring, 6 the cadmium absorber, 7 the plastic scintillator, and 8 the outer chamber.



He L C, et al. NIM-A, 2018, 880: 22-27.



Typical energy spectrum at a proton energy of 2.7 MeV



$^{103}\text{Ag}^{g+m}$ 65.7(7)min

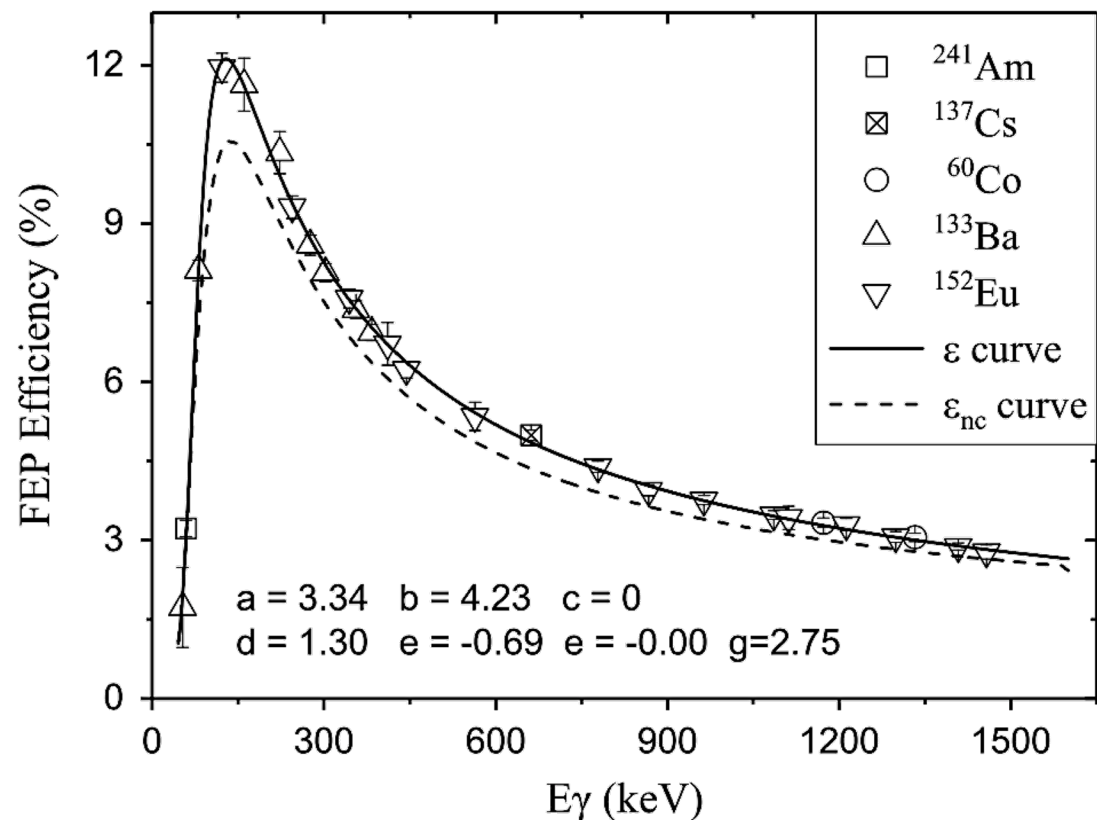
118.7 31.2(7)

148.2 28.3(5)

243.9 8.5(5)

266.9 13.4(4)

1273.8 9.4(3)

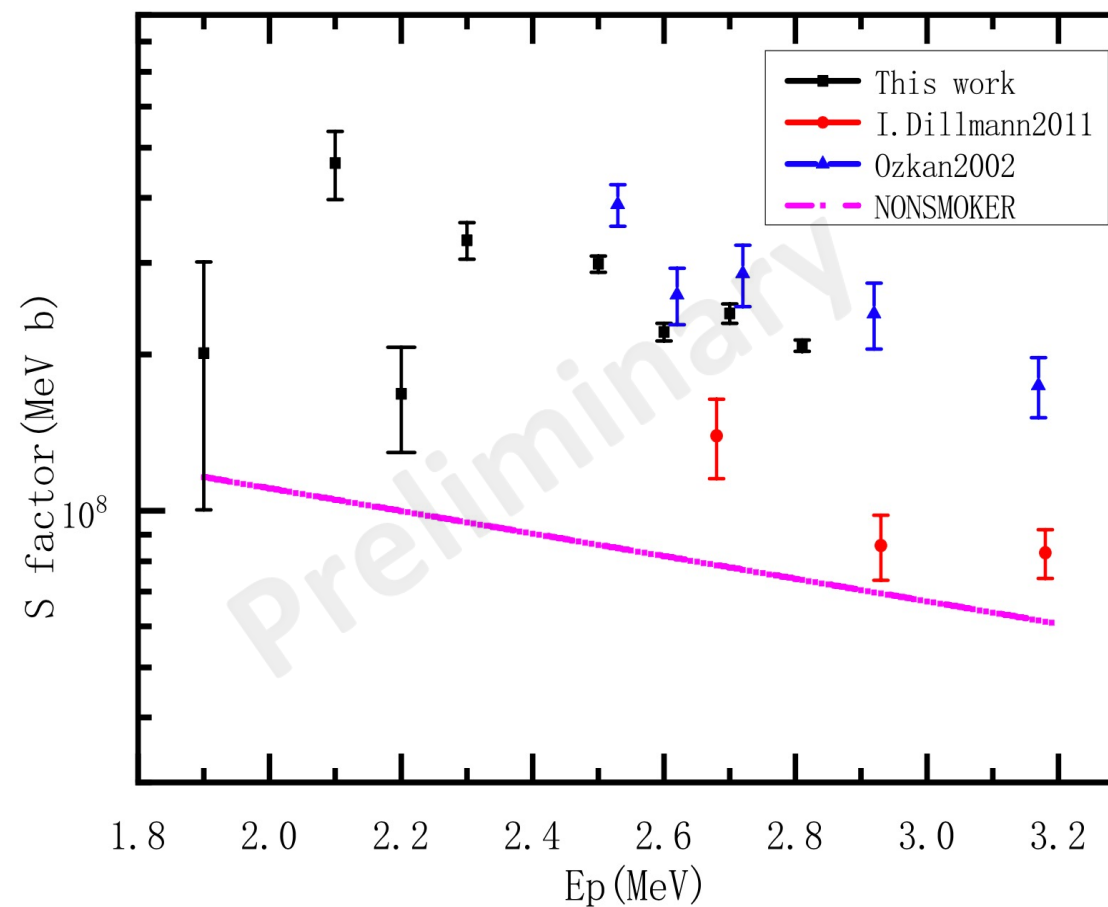


Fit Formula for Detection Efficiency Curve:

$$\ln \varepsilon_{\gamma} = a_0 + a_1 \times (\ln E) + a_2 \times (\ln E)^2 + \dots$$

$$\varepsilon_{\gamma} = \exp \left\{ \left[(a + bx + cx^2)^{-g} + (d + ey + fy^2)^{-g} \right]^{-\frac{1}{g}} \right\}$$

Dashe line: without the efficiency correction.



S-factor as a Function of Energy



Summary

1. $^{102}\text{Pd}(p,\gamma)$ is one of the important reaction associated with the abundance ^{102}Pd .
2. New cross section data of $^{102}\text{Pd}(p, \gamma)$ reaction at 1.9 MeV-3.2 MeV region had been obtained by using the activation method and the anti-Compton anti-Muon low background detector, and extended it to the lowest energy region at present.

Thanks for your attention !

C R I B